

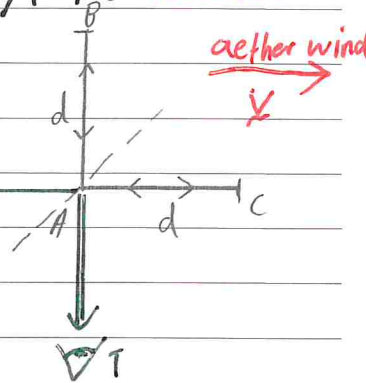
1A Special Relativity

- Classical physics proposed that light travelled through the **aether**.
- The **Michelson-Morley experiment** tested this using an interferometer.

↳ if there was aether:

$$t_{ACA} = \frac{d}{c+v} + \frac{d}{c-v}, \quad t_{ABA} = \frac{2d}{\sqrt{c^2-v^2}}$$

$$\Rightarrow \Delta t \approx \frac{dv^2}{c^3}$$



↳ but there was no shift in observed \Rightarrow no aether.

- Maxwell knew that $c = \frac{1}{\sqrt{\epsilon_0 \mu_0}}$, i.e. depends on constants unrelated to the motion of the observer.
- Einstein's postulates:
 - Speed of light (in vacuum) is the same for all observers
 - Principle of relativity**: laws of physics are the same in all inertial frames
- An important consequence is the **loss of simultaneity**: events that are simultaneous in one frame may not be in another frame.

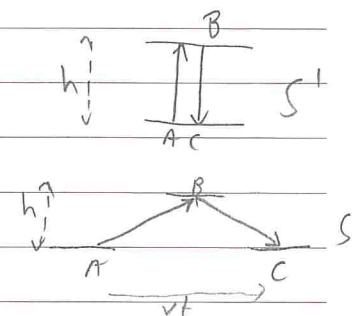
Time dilation

- Consider a light clock on a train:

- in the clock's rest frame, $\Delta t'_{AC} = \frac{2h}{c}$

- in the lab frame, $\Delta t_{AC} = \frac{2h}{\sqrt{c^2-v^2}}$

$$\Rightarrow \Delta t = \gamma \Delta t', \quad \gamma \equiv \frac{1}{\sqrt{1-v^2/c^2}}$$

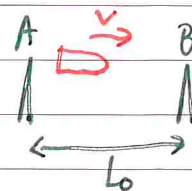


- Because $\gamma > 1$, the time interval in the moving clock is lower when viewed from a stationary frame: 'moving clocks run slow'.
- The **proper time interval** is the shortest time interval between two events, measured by a clock present at both events

- Consider two beacons separated by L_0 in their rest frame, and a space ship travelling between them.

- ship captain measures proper time $\Delta t'_{AB}$

- observer at rest measures Δt_{AB}



- The observer sees $\Delta t_{AB} = L_0/v$
- The captain sees $\Delta t'_{AB} = L/v$
- $\Rightarrow L = \frac{L_0}{\gamma}$ i.e. objects are longest in their rest frame (proper length). Length contraction otherwise
- There is no length contraction perpendicular to motion.

Evidence for SR

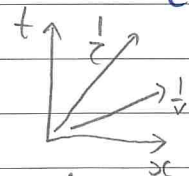
- Time-dilation in the decay of muons: Earth-based clocks measure sufficiently long time intervals for the muons to reach the surface before decaying
- Michelson-Morley: c invariant
- GPS satellites require a correction

Relativistic kinematics

- Lorentz transformation:

$$\begin{array}{l} \Delta x' = \gamma (\Delta x - v \Delta t) \\ \Delta t' = \gamma (\Delta t - \frac{v \Delta x}{c^2}) \end{array} \begin{array}{l} \xleftrightarrow{\text{Flip primes}} \\ \xleftrightarrow{\text{Flip signs}} \end{array} \begin{array}{l} \Delta x = \gamma (\Delta x' + v \Delta t') \\ \Delta t = \gamma (\Delta t' + \frac{v \Delta x'}{c^2}) \end{array}$$

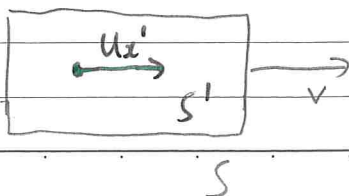
- It is often helpful to draw spacetime diagrams (one for each frame).



- While the Lorentz transforms have c as an upper bound, this does not apply to virtual objects (e.g. shadows).

Velocity addition

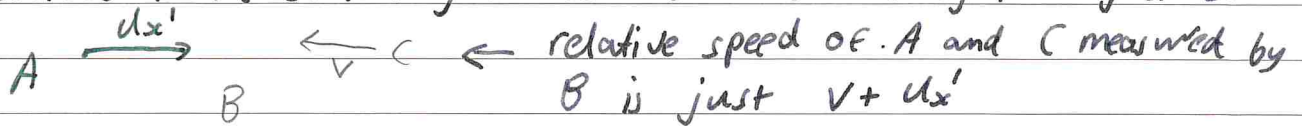
- Object moving at u_x' relative to S' , but S' moving at v relative to S



$$u_x = \frac{\Delta x}{\Delta t} = \frac{\Delta x' + v \Delta t'}{\Delta t' + v \Delta x' / c^2} \Rightarrow$$

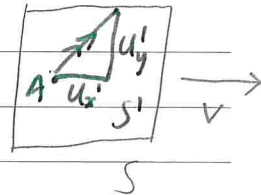
$$u_x = \frac{u_x' + v}{1 + \frac{u_x' v}{c^2}}$$

- This formula only applies when finding the velocity of the object relative to the stationary frame as measured by the object or S



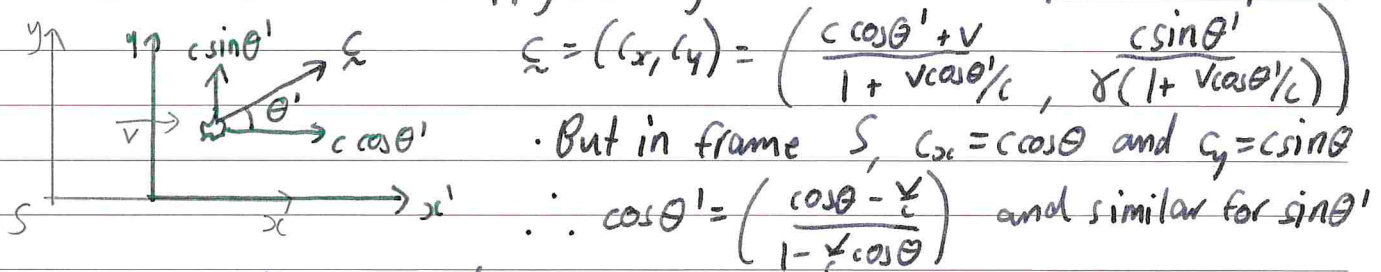
- If the motion has a transverse component, that will be affected because of time dilation

$$u_y \equiv \frac{\Delta y}{\Delta t} = \frac{\Delta y'}{\gamma(\Delta t' + v\Delta x'/c^2)} \Rightarrow u_y' = \frac{u_y}{\gamma(1 + u_x'v/c^2)}$$



Aberration of light

- It is observed that light from distant stars appears to come in at a slight angle \rightarrow originally used as evidence for aether.
- Consider a star stationary w.r.t S' , but S' moving at speed v w.r.t S . We can apply velocity addition to the photon's components:



- Thus the stellar aberration can be explained by SR.

Doppler effect

- Consider a signal with freq f' emitted by an object moving towards you at speed v . $\Delta t' = 1/f'$
- In your frame, the photons have travelled $c\Delta t'$ and the source has travelled $v\Delta t'$ by the next flash.
- \therefore photons reach your eye with $\Delta T = \frac{1}{c} (c-v)\Delta t'$

$$\Rightarrow f = f' \sqrt{\frac{1 + v/c}{1 - v/c}}$$

- Because of time dilation, there will also be a transverse Doppler shift even when the signal is perpendicular to motion.

Relativistic mechanics

- The relativistic momentum of a particle with velocity v is given by $p = \gamma v m$
- The total energy of a relativistic particle is $E = \gamma m c^2$
- The KE of a particle is the total E - rest mass E : $K = (\gamma - 1) m c^2$
- p and E conserved, but K is not.
- For a single particle:
 - $\hookrightarrow E^2 - p^2 c^2 = m^2 c^4$
 - $\hookrightarrow K^2 + 2K m c^2 = p^2 c^2$
- The E - p invariant states that the quantity $E^2 - p^2 c^2$ is constant in all inertial frames. E and p are the totals for the system, including sign in the case of p .
- Energy and momentum transform in the same way as spacetime, substituting $(ct, x, y, z) \rightarrow (E/c, p_x, p_y, p_z)$.
 - $E'/c = \gamma(E/c - v p_x/c)$
 - $p_x' = \gamma(p_x - v E/c^2)$